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1    ***In-situ* strontium and sulfur isotope investigation of the Ni-Cu-(PGE) sulfide ore-bearing Kevitsa**  
2    **intrusion, northern Finland**

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13

14   **Keywords:** strontium isotopes, sulfur isotopes, Ni-Cu-PGE sulfide ore, Kevitsa intrusion, Central  
15   Lapland greenstone belt

16

17

18

## 19    **Abstract**

20    The ~2.06 Ga Kevitsa mafic-ultramafic intrusion in northern Finland, hosts a large disseminated Ni-  
21    Cu-PGE deposit. The deposit occurs in the ultramafic olivine-pyroxene cumulates and show a  
22    range in Ni tenors varying from 4-7 wt% (regular ore) to >10 wt% (Ni-PGE ore). There are also a  
23    metal-poor sulfide mineralization (false ore) and contact mineralization that are uneconomic (Ni  
24    tenor <4 wt%).

25    The obtained  $^{87}\text{Sr}/^{86}\text{Sr}_{(i)}$  values of the Kevitsa ultramafic cumulates are highly radiogenic (>0.7045)  
26    in comparison to the estimated depleted-mantle Sr isotope ratio of ~0.702 at 2.06 Ga. The sulfur  
27     $\delta^{34}\text{S}$  values are generally higher than +2 ‰, which together with the Sr isotope data imply  
28    involvement of crustal material in the genesis of the Kevitsa intrusion and its ores. The  $^{87}\text{Sr}/^{86}\text{Sr}_{(i)}$   
29    values obtained from the ore-bearing domain of the intrusion show stratigraphic variation and  
30    exceed 0.7050, with the maximum value reaching up to 0.7109. In contrast, in rocks around the  
31    ore domain, the initial Sr isotope compositions remain more or less constant (0.7047–0.7060)  
32    throughout the intrusive stratigraphy. The isotope data suggest that the ore-bearing domain of the  
33    intrusion represents a dynamic site with multiple injections of variably contaminated magma  
34    whereas the surrounding part of the intrusion experienced a less vigorous emplacement history.  
35    No correlation is observed between the strontium and sulfur isotope compositions. This is  
36    explained by bulk assimilation of the silicate magma in a deeper staging magma chamber and  
37    variable assimilation of sulfur during magma transport into the Kevitsa magma chamber. The low  
38    level of metals in false ore and Ni-depleted nature of its olivine suggest that some sulfides may  
39    have precipitated and deposited in the feeder conduit during the initial stage of magma  
40    emplacement. Cannibalization of early-formed sulfides by later magma injections may have been  
41    important in the formation of the economic ore deposit.

## 42    **Introduction**

43

44    The evolutionary histories of mafic-ultramafic intrusive bodies may involve complex episodes of  
45    magma replenishment, magma mixing and mingling, contamination, crystal fractionation, and  
46    post-cumulus processes (e.g., DePaolo 1985; Sparks et al. 1985; Meyer and Wilson 1999; Namur et  
47    al. 2010). Whole-rock chemistry and mineral compositions are widely applied to interpret  
48    magmatic histories of intrusive bodies (e.g., Seat et al. 2007; Pang et al. 2009; Namur et al. 2010).  
49    Where these compositions change due to crystal fractionation, radiogenic or stable isotope ratios  
50    remain unaffected in closed-system processes, but may change due to addition of crustal  
51    contaminants or influx of magmas of distinct lineage into magma chambers. Consequently,  
52    isotopes have the advantage of identifying involvement of isotopically distinct magmas in the  
53    generation of igneous rock suites.

54

55    *In-situ* LA-MC-ICP-MS analysis is an effective method to determine the Sr isotope composition of  
56    plagioclase grains, potentially revealing variations in the magma composition from which the  
57    plagioclase crystallized. Grain-scale studies have been used to identify open magma chambers (Liu  
58    et al. 2014), to identify chemically distinct magmas and mixing of either magmas or minerals  
59    (Tepley et al. 1999; Seabrook et al. 2005; Yang et al. 2013a; Chen et al. 2016), and to reveal crustal  
60    contamination (Tepley and Davidson 2003). Hence, isotopes provide a tool to unravel processes  
61    operating during filling of magma chambers, including those related to sulfide ore formation in  
62    mineralized intrusions.

63 The Ni-Cu-(PGE) sulfide ore-bearing Kevitsa intrusion is one of the manifestations of the wide-  
64 spread Paleoproterozoic mafic magmatism in the Central Lapland greenstone belt, northern  
65 Finland (Hanski and Huhma 2005). Other roughly coeval magmatic sulfide mineralization is  
66 represented by the nearby Sakatti Cu-Ni-PGE deposit occurring in a small subvolcanic peridotite  
67 body (Brownscombe et al. 2015) and the komatiite-hosted Lomalampi PGE-(Ni-Cu) deposit  
68 (Törmänen et al. 2016). The Kevitsa Ni-Cu-(PGE) ore occurs in the central part of the ultramafic  
69 portion of the intrusion, whereas magmatic sulfide segregations are more commonly found at the  
70 base of differentiated mafic-ultramafic intrusions (e.g., Barnes and Lightfoot 2005). The deposit is  
71 made up of low-grade disseminated sulfides with current measured, indicated, and inferred  
72 mineral resources of 166Mt at 0.22 % Ni, 0.35 % Cu, 0.13 g/t Pt, and 0.08 g/t, Pd (data available at  
73 Boliden AB Web site). The metal content of sulfides shows an unusually large variation with their  
74 Ni tenors covering a range from ~4 wt% up to 40 wt% (Mutanen 1997; Yang et al. 2013b).

75

76 According to the interpretations by Mutanen (1997), the Kevitsa intrusion represents  
77 differentiation of a single batch of basaltic magma and in part, lithological and chemical variations  
78 reflect variable degrees of *in-situ* contamination with material from pelitic metasedimentary and  
79 mafic-ultramafic volcanogenic rocks. More recently, the injection of multiple magma pulses has  
80 been considered a more plausible explanation for the lithological and compositional variability  
81 within the ore domain and for the formation of the sulfide ores (Gregory et al. 2011; Luolavirta et  
82 al. 2018). A dynamic magma plumbing system could enable sulfide liquid to interact with a large  
83 volume of silicate magma, leading to an increase in the chalcophile element contents of the  
84 sulfides. Such an open-system behavior has been emphasized by Naldrett (1999, 2011) as one of  
85 the fundamental aspects in the formation of magmatic sulfide ores and has been regarded as

86 plausible in various well-known ore deposits, such as Voisey's Bay (Li and Naldrett 1999), Jinchuan  
87 (Song et al. 2009), Noril'sk-Talnakh (Li et al. 2003) and Uitkomst (Li et al. 2002).

88

89 Another key process in the formation of magmatic sulfide deposits is segregation of an immiscible  
90 sulfide liquid (e.g., Naldrett 2004). Sulfide saturation of mafic magma can be achieved via various  
91 mechanisms involving changes in magma compositions or P-T conditions (see Li and Ripley 2005  
92 and Ripley and Li 2013 for reviews); yet, incorporation of external sulfur is generally considered  
93 most important (e.g., Ripley and Li 2013; Keyes and Lightfoot 2010). The evidence for the presence  
94 of external sulfur is well-documented from many Ni-Cu-PGE sulfide deposits, such as Noril'sk (Li et  
95 al. 2003; Malitch et al. 2014), Voisey's Bay (Ripley et al. 1999, 2002), Jinchuan (Ripley et al. 2005;  
96 Duan et al. 2016), and Pechenga (Barnes et al. 2001), with the most convincing argument being  
97 the non-mantle-like S isotope signatures. However, some large sulfide deposits, notably Nebo-  
98 Babel (Seat et al. 2009) lack definite crustal S isotopic signatures and hence the necessity of  
99 external sulfur in generating a sulfide deposit is debatable.

100

101 In this study, we report *in-situ* Sr isotope data for plagioclase and *in-situ* S isotope data for sulfides  
102 from the Kevitsa intrusion in order to assess the nature of magma chamber processes and the role  
103 of crustal contamination in the formation of the Kevitsa intrusion and its ore deposit. Interestingly,  
104 the range of isotopic compositions and the isotope stratigraphy turned out to be markedly  
105 different in different parts of the intrusion.

106

107

## 108 **Geological setting of the Kevitsa intrusion**

109

110 The ca. 2058±4 Ma Kevitsa mafic-ultramafic intrusion is located in the Central Lapland greenstone  
111 belt (CLGB) in northern Finland (Mutanen 1997; Mutanen and Huhma 2001; Fig. 1). The CLGB is  
112 mainly composed of Paleoproterozoic komatiitic to rhyolitic metavolcanic rocks, mafic-ultramafic  
113 intrusions and sedimentary rocks with an evolutionary history ranging from ca. 2.5 Ga to 1.8 Ga.  
114 The geology of the CLGB is summarized by Hanski and Huhma (2005). A number of mafic-  
115 ultramafic intrusive bodies and volcanic rocks occur in the vicinity of the Kevitsa intrusion,  
116 including the large 2.44 Ga Koitelainen layered intrusion (Mutanen and Huhma 2001) and the Cu-  
117 Ni-PGE ore-bearing Sakatti intrusion (Brownscombe et al. 2015).

118

119 The CLGB is divided into six stratigraphic groups which are from oldest to youngest: Salla, Onkamo  
120 (currently Kuusamo Group), Sodankylä, Savukoski, Kittilä (Kittilä suite) and Kumpu Groups (Hanski  
121 and Huhma 2005; Luukas et al. 2017). The Kevitsa intrusion is hosted by the Savukoski Group  
122 volcano-sedimentary sequence consisting of interlayered phyllites, graphitic black shales and  
123 mafic to ultramafic volcanogenic rocks (Lehtonen et al. 1998; Hanski et al. 2001a; Hanski and  
124 Huhma 2005). The volcano-sedimentary country rocks are locally recrystallized to a fine-grained  
125 hornfels aureole around the intrusion. The metavolcanic rocks and, in particular, the black shales,  
126 may contain high quantities of sulfides.

127

128

129

## 130 **Kevitsa intrusion and its ore types**

131

132 The Kevitsa intrusion is composed of an approximately 1.5-km-thick ultramafic lower part and a  
133 gabbroic upper part with a minor amount of granophyre on top (Mutanen 1997; Fig. 1). The  
134 maximum thickness of the gabbroic succession at the current erosional level exceeds 500 m. In  
135 addition, dunitic rocks occur as inclusions within the Kevitsa intrusion and as a separate intrusive  
136 body (Central Dunite) in close association with the Kevitsa intrusive successions (Mutanen 1997;  
137 Yang et al. 2013b; Luolavirta et al. in press) (Fig. 1).

138

## 139 ***Rock types***

140

141 At the bottom of the Kevitsa intrusion, there is a basal series comprising pyroxenite and gabbro.  
142 The overlying ultramafic cumulates include olivine pyroxenites (OLPX), plagioclase-bearing  
143 (olivine) websterites (pOLWB) and pyroxenites (PX) (Fig. 2). The olivine pyroxenites (more  
144 precisely olivine websterites and olivine clinopyroxenites) represent the most abundant rock type  
145 and are composed of cumulus olivine (10–30 %), clinopyroxene (65–85 %), and oikocrystic  
146 orthopyroxene (0–15 %), showing ad- to mesocumulate textures (Figs. 3a, b). The pyroxenites  
147 contain less than 5 % olivine (Fig. 3c). Accessory minerals include magnetite, intercumulus  
148 plagioclase, sulfides and locally phlogopite, hornblende, ilmenite, and apatite. Plagioclase-bearing  
149 olivine websterites can be distinguished from typical olivine pyroxenites and pyroxenites by their  
150 higher contents of plagioclase (15–25 %) and orthopyroxene (15–30 %). In pOLWB, plagioclase  
151 occurs largely as an intercumulus phase defining an orthocumulate texture (Figs. 3d–e). Olivine



152 can be one of the major constituents of pOLWB (up to 15 %) or absent. Hornblende, phlogopite,  
153 magnetite, and sulfides are common accessory minerals. Fine-grained gabbros (microgabbros)  
154 with gradational contacts are found in close association with pOLWB and are considered as part of  
155 the pOLWB zones. Hence, we use the term pOLWB or pOLWB zone as a lithological unit including  
156 microgabbros.

157

158 Clinopyroxene and olivine are the prevalent constituents throughout the ultramafic zone, being  
159 generally subhedral and rather equal in size (~0.5–1.5 mm). Orthopyroxene is oikocrystic (~2–5  
160 mm), enclosing rounded to anhedral grains of olivine, clinopyroxene, and magnetite. In olivine  
161 pyroxenites and pyroxenites, plagioclase occurs as a low-mode intercumulus phase. In pOLWB,  
162 plagioclase occupies large intercumulus domains as irregularly shaped individual grains up to 5  
163 mm in size, which may enclose olivine and clinopyroxene. Aggregates of smaller  
164 anhedral/subhedral plagioclase crystals occur as well.

165

166 The Central Dunite is composed of olivine-chromite cumulates, with its modal mineralogy varying  
167 from dunite to wehrlite and feldspathic wehrlite. The dunite body shows a chemical affinity to the  
168 picritic basalts of the Savukoski Group as well as to the Kevitsa olivine-pyroxene cumulates and has  
169 been regarded as representing an initial stage of the formation of the Kevitsa intrusive suite rocks  
170 (Luolavirta et al. in press). The ore-bearing domain of the ultramafic zone is characterized by  
171 numerous rafts of dunitic rocks (up to several tens of meters in size) and komatiitic xenoliths.  
172 Pelitic xenoliths are rare within the ultramafic zone and tend to be found near the basal contact of  
173 the intrusion.

174    ***Internal stratigraphy***

175

176    In terms of lithological variation, the cumulate stratigraphy of the ultramafic rocks constituting the  
177    ore-bearing domain and the surrounding intrusion are hard to correlate. The ore domain is  
178    characterized by numerous dunitic and komatiitic inclusions, discontinuous zones of pOLWB  
179    within the OLPX (Fig. 2a) and stratigraphic fluctuations in whole-rock and mineral compositions.  
180    Around the ore domain, the rocks appear far more homogeneous and a simple stratigraphy with a  
181    typical evolutionary trend from basal series rocks via olivine pyroxenites and pyroxenites is  
182    observed (Fig. 2b). The relationship between the gabbroic zone and the ultramafic part of the  
183    Kevitsa intrusion is not well constrained. Overall, in the south, the contact appears steeply dipping  
184    and inter-fingered.

185

186    ***The deposit and ore types***

187

188    The Kevitsa Ni-Cu-(PGE) deposit is hosted by olivine pyroxenites in the central part of the  
189    ultramafic zone of the intrusion (Fig. 2a). The mineralization is made up of low-grade disseminated  
190    sulfides, with the whole-rock sulfur content typically being below 3 wt%. The low grades are  
191    compensated by the large dimensions of the ore body: it extends along the northwest axis for  
192    more than 1200 m, has a width of ~500 m and exceeds to a depth of up to 800 m. The ore body is  
193    irregular in shape and consists of several mineralized domains with variable ore grades.

194

195 Based mainly on the Ni tenor, four ore types were recognized by Mutanen (1997): “regular”, “Ni-  
196 PGE”, “false” and contact mineralization. The regular and Ni-PGE ore (Ni-Cu-(PGE) ore) comprise  
197 the economic resources, of which the regular ore type covers ~95 % by volume (Santaguida et al.  
198 2015). The regular ore type generally has a Ni tenor in the range of 4–7 %, with Ni/Cu ratio falling  
199 below 1 and the PGE content varying from “low” (~100 ppb of Pt) to high (~750 ppb of Pt). The Ni-  
200 PGE ore occurs as lens-like discontinuous bodies and is characterized by a high Ni tenor of >10 %,  
201 low copper and high PGE contents (400 to 3000 ppb of Pt). In the preliminary characterization of  
202 the ore types by Mutanen (1997), the term “transitional ore” was also used to an ore type with an  
203 intermediate composition between the regular and Ni-PGE ore. However, based on the chemical  
204 affinity of the transitional ore towards the Ni-PGE ore (Hanski et al. 1997, discussed below), the  
205 transitional ore can be considered lower-grade Ni-PGE ore. The uneconomic mineralization that is  
206 called false ore has a low Ni tenor (0.5–4 wt%) and low Pd and Pt contents, ranging from few tens  
207 of ppb to values below the detection limit of 10 ppb. The contact mineralization can locally  
208 comprise semi-massive sulfides but shows very low metal contents (Ni tenor 1–2 wt%). It is worth  
209 emphasizing that there exists a continuous range of ore compositions with two broad  
210 compositional trends from the regular ore: one towards ores very rich in Ni (and PGE) and low in  
211 Cu and the other towards mineralized rocks almost totally devoid of Ni, Cu and PGE.

212

213 There is a clear relationship between the metal tenors of the different ore types and their sulfide  
214 mineral assemblages. The dominant ore minerals are pentlandite and chalcopyrite in the regular  
215 ore type and pentlandite, millerite, and heazlewoodite in the Ni-PGE ore type. The false ore and  
216 contact mineralization are dominated by pyrrhotite. The sulfides occur together with a small  
217 amount of magnetite in the interstitial spaces between olivine and pyroxene grains. Besides the

218 well-developed magmatic textures of the sulfides, the magmatic origin of the mineralization is  
219 reflected, for example, in positive correlations between nickel, copper and sulfur contents as well  
220 as between platinum and palladium concentrations (not shown). Pd/Pt ratios for false ore, regular  
221 ore and Ni-PGE ore are similar ( $\sim 0.6$ ; Le Vaillant et al. 2016). Effects of hydrothermal alteration to  
222 the Kevitsa sulfide ore has recently been discussed by Le Vaillant et al. (2016) who argue that of  
223 the base metals, copper (and Au) may have been locally redistributed, but to what extent is not  
224 well established.

225

226 Besides metal tenors, the ore types show differences in their REE contents and isotope  
227 compositions. The most intriguing mineralization type is the Ni-PGE ore, which has abnormally  
228 high contents of Ni, not only in the sulfides, but also in the primary silicates (olivine and  
229 clinopyroxene), being in this sense very “primitive” (Mutanen 1997; Yang et al. 2013b; Luolavirta  
230 et al. 2018). Yet, this ore type has significantly LREE-enriched chondrite-normalized REE patterns  
231 ( $Ce_N/Yb_N \sim 7$ , Hanski et al. 1997; Luolavirta et al. 2018) and records low initial  $\epsilon_{Nd}$  ( $-6.4$ ; Huhma et  
232 al. 2017 in review) (Fig. 4). Furthermore, Luolavirta et al. (2018) demonstrated that there is a  
233 mineralogical difference between the ore types, as the host rocks to the Ni-PGE ore type tend to  
234 be virtually devoid of orthopyroxene (olivine clinopyroxenites). The  $\epsilon_{Nd}$  values for regular and false  
235 ore are similar ( $-3.4$ ; Huhma et al. 2017 in review) (Fig. 4) with both of them showing mildly LREE-  
236 enriched REE patterns ( $Ce_N/Yb_N \sim 2$ ; Hanski et al. 1997; Luolavirta et al. 2018), but they differ in  
237 terms of their S isotope compositions. Average whole-rock  $\delta^{34}S$  values of  $+3.8$ ,  $+6.1$  and  $+8.2\%$   
238 have been reported for regular, Ni-PGE (including the transitional ore) and false ore, respectively  
239 (Grinenko et al. 2003) (Fig. 5).

240

## 241 **Sampling and analytical methods**

### 242 ***Samples***

243

244 Samples for *in-situ* strontium and sulfur isotope analyses were collected from three drill cores to  
245 cover the stratigraphy through the ultramafic cumulate succession and to obtain isotope data  
246 from the different ore types. Drill core KV-103 represents the ore domain, being relatively deep  
247 (~750 m) and intersecting both the regular and Ni-PGE ore types. Drill cores KV-280 and KV-297  
248 are located a few hundred meters outside the ore-bearing domain and intersect a false ore-type  
249 mineralization and contact mineralization. The compositions of major minerals (olivine,  
250 clinopyroxene, orthopyroxene) of these drill cores are discussed in another paper (Luolavirta et al.  
251 2018).

252

### 253 ***Analytical methods***

254

255 *In-situ* Sr isotope analyses of plagioclase were performed by laser ablation multi-collector  
256 inductively coupled plasma mass spectrometry (LA-MC-ICP-MS) using a Nu Plasma HR mass  
257 spectrometer and a Photon Machine Analyte G2 laser microprobe at the Geological Survey of  
258 Finland in Espoo. Samples were ablated in He gas (gas flows = 0.4 and 0.1 l/min) within a HelEx  
259 ablation cell (Müller et al. 2009). Strontium isotope analyses of plagioclase were made in static  
260 ablation mode employing the following parameters: beam diameter 110 µm, pulse frequency 10  
261 Hz, beam fluence 2.07 J/cm<sup>2</sup>. The MC-ICP-MS instrument was equipped with 9 Faraday detectors  
262 and amplifiers with 10<sup>11</sup> Ω resistors. During the laser ablation, data were collected in static mode

263 ( $^{84}\text{Sr}$ - $\text{Kr}$ ,  $^{85}\text{Rb}$ ,  $^{86}\text{Sr}$ - $\text{Kr}$ ,  $^{87}\text{Rb}$ - $\text{Sr}$ ,  $^{88}\text{Sr}$ ). Measured isotope ratios were corrected for instrument  
264 fractionation applying an exponential law and a  $^{86}\text{Sr}/^{88}\text{Sr}$  value of 0.1194. The isobaric interference  
265 of  $^{87}\text{Rb}$  on  $^{87}\text{Sr}$  was monitored and corrected using the  $^{85}\text{Rb}$  ion signal and a value of 0.38571 for  
266 the  $^{87}\text{Rb}/^{85}\text{Rb}$  ratio. The isobaric interference of  $^{86}\text{Kr}$  on  $^{86}\text{Sr}$  was corrected using a 30 s background  
267 measurement, preceding every ablation. Strontium isotope ratios were age-corrected to 2058 Ma  
268 based on  $^{85}\text{Rb}/^{86}\text{Sr}$  ratios. The average total Sr signal obtained for plagioclase samples was 0.4 V.  
269 Under these conditions, 120 s of ablation are needed to obtain an internal precision of  $\leq \pm 0.00007$   
270 ( $1\sigma$ ). The decay constant of  $^{87}\text{Rb}$  of  $1.3968 \times 10^{-11} \text{ y}^{-1}$  given by Rotenberg et al. (2012) was used in  
271 all calculations. The accuracy of the laser ablation protocol was verified throughout the day of  
272 measurement by repeated analysis of an in-house plagioclase standard from a megacryst of the  
273 Cameroon volcanic chain (sample Mir a; Rankenburg et al. 2004). The laser ablation parameters  
274 were similar for the samples and standard. During the course of this study, the measured  $^{87}\text{Sr}/^{86}\text{Sr}$   
275 values for the standard ranged from 0.70300 to 0.70319 and averaged  $0.70310 \pm 0.00008$  ( $2\sigma$ ,  
276  $n=30$ ), which is similar to the TIMS value of  $0.70311 \pm 0.0001$  ( $2\sigma$ ) obtained by Rankenburg et al.  
277 (2004). Variations in the  $^{87}\text{Sr}/^{86}\text{Sr}$  ratio of the plagioclase standard over the course of the study are  
278 given in Electronic Supplementary Material 1 (ESM 1).

279

280 *In-situ* sulfur isotopes analyses of pyrrhotite and pyrite were performed employing the same  
281 instrument as for the Sr isotope analyses. Samples were ablated in He gas (gas flows = 0.4 and 0.1  
282 l/min) within a HelEx ablation cell (Müller et al. 2009). Sulfur isotopes were measured at medium  
283 resolution. During the ablation, data were collected in static mode ( $^{32}\text{S}$ ,  $^{34}\text{S}$ ). Sulfide samples were  
284 ablated at a spatial resolution of 50 micrometers, using a laser beam fluence of  $0.83 \text{ J/cm}^2$  at 5 Hz.  
285 The total S signal was typically 0.5–4.0 V. Under these conditions, after a 20 s baseline

286 measurement, 30–40 s of ablation is needed to obtain an internal precision in  $^{34}\text{S}/^{32}\text{S}$  of  $\leq \pm$   
287 0.000005 (1 SE). In-house pyrrhotite and pyrite standards were utilized for external standard  
288 bracketing and quality control of analyses. The standards used have been analyzed by gas mass  
289 spectrometry. For the quality control pyrite standard, our measured  $\delta^{34}\text{S}_{\text{CDT}}$  (‰) value is  $+6.61 \pm$   
290 0.45 (n=29) against the gas mass spectrometer-determined value of  $+5.8 \pm 0.3$  (‰). For pyrrhotite,  
291 these values are  $4.4 \pm 0.3$  (‰,  $1\sigma$ , n=13) and  $+4.8 \pm 0.3$  (‰,  $1\sigma$ , n=3), respectively.

292

293 Plagioclase compositions were determined using a JEOL JXA-8200 electron microprobe at the  
294 Center of Microscopy and Nanotechnology, the University of Oulu, with a standard built-in ZAF  
295 correction routine. The analytical conditions were an accelerating voltage of 15 kV and a beam  
296 current of 30 nA. Peak counting times on major elements were 60 s and 30 s for the background.  
297 The whole-rock trace-elements used in this study were provided by the mining company. Whole-  
298 rock major and trace element abundances were produced at an ALS geochemistry laboratory using  
299 ICP-MS/ICP-AES after four-acid near-total sample digestion and precious metals by ICP-MS/ICP-  
300 AES after fire assay pre-concentration.

301

## 302 **Results**

### 303 ***Sr isotope composition of plagioclase***

#### 304 ***General features***

305 Only fresh and nonfractured interstitial plagioclase grains were selected for laser ablation to  
306 minimize alteration effects to the Rb-Sr isotope system. The Rb/Sr ratios of the analyzed  
307 plagioclase grains are low and relatively constant ( $\leq 0.005$ ) indicating very little growth of

308 radiogenic  $^{87}\text{Sr}$  by  $^{87}\text{Rb}$  decay since the crystallization of the Kevitsa intrusion at 2.06 Ga. As stated  
309 above, results of repeated analyses of in-house standard are within the recommended value,  
310 verifying the accuracy of our laser ablation protocol (see ESM 1). Representative isotope data with  
311  $2\sigma$  error margins and anorthite contents of plagioclase are presented in Table 1 and the whole  
312 dataset is tabulated in Electronic Supplementary Material 2 (ESM 2).

313

314 It must be noted that the low mode of plagioclase restricted the number of spot analyses in OLPX  
315 and PX samples, whereas for pOLWB, a more comprehensive sampling could be conducted.  
316 Nevertheless, the measured  $^{87}\text{Sr}/^{86}\text{Sr}_{(i)}$  ratios are heterogeneous in each sample (Fig. 3) and  
317 overall, no observable correlations with An contents and  $^{87}\text{Sr}/^{86}\text{Sr}_{(i)}$  could be established.  
318 Relationships between the anorthite content and  $^{87}\text{Sr}/^{86}\text{Sr}_{(i)}$  in plagioclase are illustrated in  
319 Electronic Supplementary Material 3 (ESM 3).

320

#### 321 ***Drill cores KV297 and KV280 (outside the ore domain)***

322

323 The initial  $^{87}\text{Sr}/^{86}\text{Sr}$  ratios in the samples collected from the "unmineralized" part of the Kevitsa  
324 intrusion vary from 0.7047 to 0.7060 (Figs. 6 and 7). In one sample of false ore (KV280 ~520 m)  
325 and in the gabbro at the top of drill core KV-297, individual spot analyses show slightly higher  
326 initial ratios (0.7068 and 0.7065, respectively). However, the observed range of Sr isotope  
327 compositions remain relatively constant throughout the stratigraphy.

328



329 Overall, the An contents increase slightly from the base upwards, after which they remain  
330 constant in OLPX and then decrease towards the top of the ultramafic unit. One sample (at a  
331 depth of around 750 m in drill core KV280) stands out in having markedly lower An contents. This  
332 sample contains abundant primary phlogopite, which is not common in other studied samples.  
333 The only cumulus plagioclase analyzed in this study is from the gabbro sample at the top of drill  
334 core KV-297. This sample shows a highly heterogeneous plagioclase composition but the Sr  
335 isotope ratios are in the same range as in the underlying ultramafic cumulates. It appears that the  
336 cores of individual plagioclase grains record the lowest  $^{87}\text{Sr}/^{86}\text{Sr}_{(i)}$  ratios ( $\sim 0.7055$ ) in this gabbroic  
337 sample (see ESM 3).

338

#### 339 ***Drill core KV103 (ore domain)***

340

341 In drill core KV103, representing the ore-bearing domain of the ultramafic zone,  $^{87}\text{Sr}/^{86}\text{Sr}_{(i)}$  varies  
342 considerably from 0.7050 to as high as 0.7109 (Fig. 8). The highest Sr isotope ratios, 0.7089–  
343 0.7109, were measured from the host rocks to the Ni-PGE ore type, which is consistent with the  
344 highly non-radiogenic Nd isotope composition of this ore type (Hanski et al. 1997, Fig. 4). Also a  
345 marked peak in La/Nb coincides with the Ni-PGE ore, indicating elevated LREE contents in the host  
346 rocks. The mineral compositions and Sr isotope ratios vary widely in the olivine pyroxenites below  
347 the pOLWB. The OPLXs above the pOLWB unit are characterized by more uniform Sr isotope ratios  
348 of 0.7055 to 0.7073 and a smooth decrease in the An contents of plagioclase (except for a sharp  
349 increase in the uppermost portion of the drill core), but oscillatory variations in the contents of  
350 trace elements, such as Zr, and in the olivine compositions. Overall, the  $^{87}\text{Sr}/^{86}\text{Sr}_{(i)}$  values are  
351 somewhat higher in comparison to the values obtained for rocks outside the ore domain ( $\sim 0.7047$

352 to 0.7060). The samples of regular ore yielded initial Sr isotope ratios of 0.7055 to 0.7068, similar  
353 to those of the false ore and consistent with the overlap in the initial  $\epsilon_{\text{Nd}}$  values between these ore  
354 types (Hanski et al. 1997, Fig. 4).

355

356 The  $^{87}\text{Sr}/^{86}\text{Sr}_{(\text{i})}$  values within the pOLWB zone show a progressive upward increase, which  
357 correlates with a decrease in the An contents of plagioclase and an increase in the Fo contents of  
358 olivine. Sample-scale and intra-grain variations in  $^{87}\text{Sr}/^{86}\text{Sr}_{(\text{i})}$  are large (Figs. 3d–f, ESM 3). Core to  
359 rim traverses across selected plagioclase grains did not reveal any systematic variations in the An  
360 content and/or Sr isotope ratio. One sample is an exception where the core domains of individual  
361 grains tend to record the lowest  $^{87}\text{Sr}/^{86}\text{Sr}_{(\text{i})}$  values (KV-103-541.14, see ESM 3) and in this case, it  
362 can be postulated that the plagioclase cores represent an early cumulus phase mantled by  
363 overgrowths with a variable isotope signature. The microgabbros seem to record two groups of  
364 plagioclase with different  $^{87}\text{Sr}/^{86}\text{Sr}_{(\text{i})}$ . However, the small number of analyses may have generated  
365 an analytical bias, since distinct ratios were measured from single grain and no correlation  
366 between the  $^{87}\text{Sr}/^{86}\text{Sr}_{(\text{i})}$  ratio and An content of that spot can be demonstrated.

367

368 A sharp decrease in the Sr isotope ratio at the upper pOLWB–OLPX contact coincides with a minor  
369 increase in the olivine Fo contents and the occurrence of ore grades of Ni-Cu sulfides (regular ore).

370

371

372 ***In-situ S isotope analyses***

373

374 Sulfur isotope compositions were analyzed for pyrrhotite and, in two samples, also for pyrite. In  
375 addition to the different ore types,  $\delta^{34}\text{S}$  values were also measured for sulfides from "barren"  
376 rocks with little sulfides. Representative S isotope analyses are listed in Table 1. The whole dataset  
377 is provided in Electronic Supplementary Material 4 (ESM 4) and illustrated in Fig. 5.

378

379 ***Drill core KV297 and KV280 (outside the ore domain)***

380

381 The false ore samples from drill core KV-280 (outside the ore domain) record a relatively heavy  
382 sulfur isotope compositions, with  $\delta^{34}\text{S}$  ranging from +5.1 to +8.0 ‰ and the median  $\delta^{34}\text{S}$  value  
383 being ~+6.5 ‰ (Figs. 5 and 7). For false ore samples, Grinenko et al. (2003) report a wide range of  
384 whole-rock  $\delta^{34}\text{S}$  values from +3.4 up to +18.6 ‰. Most of these whole-rock S isotope data cluster  
385 in the range from +5.5 to +11.7 ‰, being broadly in line with the measured *in-situ* values of this  
386 study (Fig. 5). The  $\delta^{34}\text{S}$  values in S-poor samples from drill cores KV297 and KV280 range between  
387 +2.5 and +6, being generally  $\geq +3$  ‰ in the lower parts of these drill cores and  $\leq 3$  ‰ higher in the  
388 stratigraphy (Figs. 6 and 7).

389

390 ***Drill core KV103 (within the ore domain)***

391

392 The obtained  $\delta^{34}\text{S}$  values for the regular ore type vary from +3.6 to +5.5 ‰, with a median of ~+4  
393 ‰, consistent with the previously measured whole-rock  $\delta^{34}\text{S}$  values of +2.0 to +4.6 ‰ (+3.8 ‰ on

average; Grinenko et al. 2003, Fig. 5). The samples of the Ni-PGE ore type record  $\delta^{34}\text{S}$  values of +1.3 to +4.0 ‰ with a median value of +2.6 ‰ (Fig. 5). Interestingly, the obtained *in-situ*  $\delta^{34}\text{S}$  values for Ni-PGE ore are lower than the whole-rock values of +3.0 to +8.8 ‰ (+6.1 ‰ on average) reported by Grinenko et al. (2003) (Fig. 5).

398

The S isotope compositions remain constant in the OLPX below the pOLWB whereas the mineral compositions and Sr isotope ratios vary (Fig. 8). In pOLWB, S isotopic compositions are heavy and vary considerably from ~+4‰ up to ~+12‰ with stratigraphic fluctuations. The S isotope compositions show mantle-like (from -0.4‰ to +1.55‰) values at a depth of around 300 m, which coincide with a reversal in the whole-rock Zr contents and is roughly coincidental with the onset of a reversal in olivine composition observed few tens of meters below this depth (see Fig. 8). No correlation exists between the S and Sr isotope compositions.

406

## 407 Discussion

### 408 *Petrogenetic implications for the formation of the Kevitsa ultramafic cumulates*

409

Stratigraphic variations and reversals in mineral compositions and whole-rock element abundances and/or changes in the types and proportions of fractionating minerals are generally considered to indicate open magma chamber processes and periods of magma recharge in igneous rock suites (e.g., Cox and Hawkesworth 1985; Eales et al. 1986, 1990). Radiogenic isotopes (such as Sr) may provide further evidence for such magma chamber processes (e.g., Namur et al. 2010; Liu et al. 2014) by identifying influxes of isotopically different magmas.

416 This study has shown that compared to the estimated 2.06 Ga depleted-mantle or bulk-Earth Sr  
417 isotope ratios of 0.7018 and 0.7023, respectively (O’Nions et al. 1979), the calculated  $^{87}\text{Sr}/^{86}\text{Sr}_{(i)}$   
418 values of plagioclase ( $>0.7045$ ) are highly radiogenic throughout the Kevitsa ultramafic cumulates,  
419 implying strong involvement of crustal material in their genesis. Similarly, the  $\delta^{34}\text{S}$  values  
420 measured for pyrrhotite and pyrite vary from -0.42 to +8.00 ‰, being generally greater than +2 ‰  
421 (Fig. 5) and hence heavier than uncontaminated mantle-derived sulfur, which has been estimated  
422 to have  $\delta^{34}\text{S}$  values from -2 to +2 ‰ (Ripley and Li 2003).

423

424 The Sr isotope ratios of the Ni-PGE ore (0.709–0.711) are generally much higher than those  
425 ( $\sim 0.704$  to 0.709) obtained by whole-rock (Eales et al. 1990; Lee and Butcher 1990; Kruger 1994;  
426 Mitchell et al. 1998) or plagioclase analyses (Seabrook et al. 2005; Yang et al. 2013a; Mangwegape  
427 et al. 2016; Wilson et al. 2017) of Lower and Main zone rocks from the Bushveld Complex, which is  
428 similar in age to the Kevitsa intrusion (Fig. 4). There are no Sr isotope data on the immediate  
429 country rocks or any other rock type from the vicinity of the Kevitsa intrusion to compare with, but  
430 Kröner et al. (1981) have published Sr isotope data from the 3.1 Ga Tojottamanselkä gneiss dome  
431 located around 10 km north of Kevitsa. At 2.06 Ga, these gneisses had  $^{87}\text{Sr}/^{86}\text{Sr}$  ratios of 0.707–  
432 0.711, which overlap the upper part of the  $^{87}\text{Sr}/^{86}\text{Sr}_{(i)}$  range measured for Kevitsa rocks. The  
433 similarity of the  $^{87}\text{Sr}/^{86}\text{Sr}_{(i)}$  values suggests that the contaminant end-member must have had even  
434 more radiogenic Sr than that of the ancient gneisses of the Tojottamanselkä dome.

435

436 It is worth emphasizing that the Sr isotopes were analyzed using intercumulus plagioclase. It is  
437 well-known that the interstitial liquid may percolate within the cumulus pile and hence its  
438 composition may not directly reflect that of the coexisting cumulus phases (e.g. Boudreau and

439 McCallum 1992; Karykowski and Maier 2017). However, in the case of the Ni-PGE ore, the peculiar  
440 features of the silicates (enrichment in LREE, highly radiogenic Sr in plagioclase, highly negative  
441 initial whole-rock  $\epsilon_{\text{Nd}}$ , and Ni-rich olivine) as well as the sulfide phase (high Ni tenor) seem to be  
442 characteristic for all analyzed samples. In order to preserve these unique features, neither the  
443 sulfide liquid nor the intercumulus silicate melt could have migrated extensively with respect to  
444 each other or with respect to the cumulus minerals.

445

#### 446 ***Formation of the "unmineralized" domain of the intrusion***

447

448 The ultramafic cumulates around the ore domain (drill cores KV-297 and KV-280) show modest  
449 fluctuations in whole-rock Zr and olivine Fo contents, rather constant La/Nb ratio, and record  
450 predictable fractionation trends from the basal pyroxene-gabbro via olivine pyroxenites to  
451 pyroxenites (and gabbro in drill core KV-297) (Figs. 6 and 7). Also, the range of the Sr isotope  
452 compositions remain constant throughout the stratigraphy. From the base upwards, reverse  
453 fractionation trends can be observed, particularly in the Zr contents. Such marginal reversals are a  
454 common feature for mafic intrusive bodies worldwide and are considered to reflect prolonged  
455 magma emplacement (Latypov et al. 2011; Egorova and Latypov 2012a, 2012b).

456

457 The recent model by Le Vaillant et al. (2017) proposes that the metal-poor sulfide dissemination  
458 (false ore) around the ore domain formed within early-stage xenolith-laden sill-like intrusions in  
459 which high viscosity circumstances restricted mixing of the magma and sulfide liquid, resulting in  
460 low metal tenors. This model is not consistent with the observed homogeneity in the chemical and

461 Sr isotopic compositions of the ultramafic cumulates around the ore domain. We propose that the  
462 rocks around the ore domain formed by continuous inputs of chemically and isotopically rather  
463 uniform basaltic magma into the Kevitsa magma chamber, followed by crystal fractionation.

464

465 The S isotope compositions in false ores are markedly heavier in comparison to "barren" rocks (S  
466 <0.5 wt%). Also the chalcophile metal contents, particularly those of Pt and Pd, as well as the Ni  
467 content in olivine (discussed further below) are relatively low. Given the predictable evolutionary  
468 paths and compositional homogeneity of several hundred meters thick succession of olivine  
469 pyroxenite in the drill cores located outside the ore domain, the inputs of magma hardly took  
470 place as discrete periods of magma emplacement, yet the metal content and the sulfide budget of  
471 the inflowing magma batches varied. Notably, the Sr isotopic compositions remain rather  
472 homogeneous throughout the stratigraphy whereas the S isotopic compositions vary. This is best  
473 explained by bulk contamination of the magma at some deep-seated staging chamber and variable  
474 degrees of assimilation of crustal sulfur during transportation of the magma into the Kevitsa  
475 magma chamber.

476

#### 477 ***Formation of the ore-bearing domain of the intrusion***

478

479 The compositional stratigraphy of drill core KV-103 is characterized by obvious fluctuations in the  
480 whole-rock Zr contents and olivine compositions and variations in the Sr and S isotope signatures  
481 (Fig. 8). Luolavirta et al. (2018) suggest that the stratigraphic variations in mineral and whole-rock  
482 compositions in the drill cores from the ore domain reflect episodes of magma replenishment. The

483 high  $^{87}\text{Sr}/^{86}\text{Sr}_{(i)}$  values (0.709–0.711) in the host rocks to the Ni-PGE ore and their marked decrease  
484 at the level of the regular ore further support open magma chamber processes, yet restricted to  
485 the ore-bearing domain of the intrusion. The sulfur isotope ratios in drill core KV-103 show no  
486 correlation with the Sr isotope compositions, indicating that the magma pulses assimilated  
487 variable degrees of silicate components and sulfur from the country rocks or the isotopic  
488 signatures were generated via separate contamination processes (as proposed based on isotopic  
489 data outside the deposit area in drill cores KV-297 and KV-280).

490

491 ***Sample-scale Sr isotope heterogeneity and intra-mineral disequilibrium in pOLWB***

492

493 Plagioclase-bearing olivine websterites are only found in the ore domain where they form irregular  
494 zones that locally act as marker horizon for regular ore-type sulfides (as in drill core KV-103).  
495 According to the model by Mutanen (1997), whereby all lithological and chemical variations are a  
496 result of variable degrees of *in-situ* contamination, the zones of pOLWB could be interpreted to  
497 reflect significant incorporation of pelitic material in their genesis. More recently, the generation  
498 of the zones of pOLWB has been related to fractionation of individual magma pulses, with sulfide  
499 mineralization occurring near the basal parts of these pulses (Gregory et al. 2011). The obtained Sr  
500 isotope ratios show a progressive increase up-hole through the pOLWB, which correlates with a  
501 decreasing anorthite content of plagioclase and an increasing Fo content of olivine. Such mineral  
502 compositional and isotopic profiles cannot be explained by simple crystal fractionation.

503



504 In the pOLWB zones, the sample-scale variations in  $^{87}\text{Sr}/^{86}\text{Sr}_{(i)}$  are relatively large and core-to-rim  
505 traverses reveal considerable intra-mineral isotopic differences (Fig. 3d-f, ESM3). Grain-scale  
506 isotopic disequilibrium has been explained by interaction of the initial magma with a new,  
507 isotopically distinct magma influx (Tepley et al. 2000; Gao et al. 2015), by late-stage infiltration of  
508 isotopically distinct residual melt or fluid through the cumulate pile and its reaction with solid  
509 crystals in the pile (Chutas et al. 2012; Yang et al. 2013a), or crystallization accompanied by  
510 contamination (Templey and Davidson 2003). These processes should result in systematic core-to-  
511 rim zonation in isotope compositions of individual grains. Chutas et al. (2012) also observed  
512 different Sr isotope compositions between large and small orthopyroxene grains in the Lower  
513 Zone of Bushveld Complex and proposed that the larger grains grew at the expense of smaller  
514 ones in the presence of fluid with different  $^{87}\text{Sr}/^{86}\text{Sr}_{(i)}$ .

515

516 To explain the wide sample- and grain-scale spread of  $^{87}\text{Sr}/^{86}\text{Sr}_{(i)}$  in pOLWB is not straightforward.  
517 First of all, we have not observed either clear trends in  $^{87}\text{Sr}/^{86}\text{Sr}_{(i)}$  values between plagioclase cores  
518 and rims or any distinct textural populations that record more homogeneous  $^{87}\text{Sr}/^{86}\text{Sr}_{(i)}$  values.  
519 Also, no apparent correlation exists between  $^{87}\text{Sr}/^{86}\text{Sr}_{(i)}$  and An, suggesting that the plagioclase  
520 crystallized from an isotopically heterogeneous intercumulus liquid. However, it must be noted  
521 that the analyzed plagioclase grains are irregular in shape and the actual position of core domains  
522 (assumed early cumulate) cannot be accurately defined. The  $^{87}\text{Sr}/^{86}\text{Sr}_{(i)}$  values in the pOLWB zone  
523 approach those measured from the underlying OLPX unit hosting Ni-PGE ore and hence, upwards  
524 percolation of residual melt originating from the underlying OLPX could be a viable explanation for  
525 the radiogenic Sr in the pOLWB. However, in this case, it would be reasonable to expect an  
526 opposite sense of isotopic change, i.e., an upwards decrease in  $^{87}\text{Sr}/^{86}\text{Sr}_{(i)}$  in the pOLWB. A similar

527 assumption would apply to the interaction (mixing/mingling or melt percolation) between pOLWB  
528 with the overlying OLPX.

529

530 A concomitant up-hole increase in Mg# of orthopyroxene and initial Sr isotope ratio, similar to that  
531 observed in pOLWB zone, has been described from the basal unit of the Lower Main Zone of the  
532 Bushveld Complex (Mitchell 1990; Mitchell et al. 1998). It was explained by mixing of resident  
533 residual magma in the chamber with new influxes of the Main Zone-type, high- $^{87}\text{Sr}/^{86}\text{Sr}_{(i)}$  magma.  
534 Progressive mixing or mingling of unsolidified pOLWB and the overlying, more primitive OLPX  
535 could potentially explain the reverse fractionation trends in mafic minerals, yet as discussed  
536 above, the Sr isotopic profile across the pOLWB - OLPX contact does not indicate any significant  
537 interaction between these rock units. It must be noted that the compositions of ferromagnesian  
538 minerals may have equilibrated with variable degrees of trapped liquid resulting in modifications  
539 in their primary compositions (Barnes 1986). This could potentially explain the up-hole increase in  
540 the olivine Fo contents in the pOLWB zone. Petrography or whole-rock compositions of the  
541 studied samples, however, do not reveal any significant differences in the amount of trapped  
542 liquid.

543

544 In the Rum layered intrusion, Templey and Davidson (2003) observed a trend of upward-increasing  
545 plagioclase  $^{87}\text{Sr}/^{86}\text{Sr}_{(i)}$  towards a lithological unit contact, accompanied by isotopic disequilibrium  
546 between cores and rims of some grains. The authors propose that the magma from which the  
547 plagioclase crystallized underwent progressively larger degrees of *in-situ* contamination. The  
548 model predicts that isotopic disequilibrium develops in minerals initially growing near the  
549 roof/margins of the intrusion where crustal contamination can be assumed to be most effective.

550 The presence of microgabbros within the zones of pOLWB supports the crystallization of the  
551 pOLWB in proximity to the wall rocks. Similar considerations can be inferred from the relatively  
552 large variation in the S isotope composition in the pOLWB zones. However, we cannot  
553 unambiguously state whether the pOLWB potentially represents a roof sequence of a discrete  
554 magma pulse or a separate intrusive phase(s) at the time when the geometry of the Kevitsa  
555 intrusion was different from what is currently observed, or autolith(s) of some former marginal-  
556 phase rock of the Kevitsa intrusion.

557

#### 558 ***Implications for ore-forming processes***

#### 559 ***Isotopic constraints on the origin of the Ni-Cu-(PGE) ore***

560

561 Many authors have discussed and reviewed the theoretical aspects related to the formation of  
562 magmatic Ni-Cu-PGE deposits (e.g., Naldrett 1999, 2004, 2011; Maier et al. 2001; Arndt et al.  
563 2005; Barnes and Lightfoot 2005; Maier and Groves 2011; Song et al. 2011). Briefly, the key factors  
564 include: i) a reasonably high degree of mantle melting generating a parental mafic-ultramafic  
565 magma with adequate concentrations of metals, ii) emplacement of the magma into or onto the  
566 crust with minimum prior fractionation of olivine or sulfides, iii) contamination of the magma with  
567 crustal materials, promoting sulfide saturation, iv) interaction of sulfides with a large volume of  
568 magma resulting in enrichment of the sulfides in metals and v) mechanical concentration of  
569 sulfides to economic levels.

570

571 The solubility of sulfide in a mafic magma is known to increase with falling pressure and hence  
572 magmas ascending to shallow crustal depths are likely sulfur undersaturated (e.g., Mavrogenes &  
573 O'Neill 1999). Consequently, a process is required to bring the magma to sulfur saturation under  
574 low-pressure conditions. In the case of magmatic Ni-Cu sulfide deposits, this is generally achieved  
575 via various processes of contamination, such as addition of silica or volatiles (see Ripley and Li  
576 2013 for a review). Evidently, the most feasible process is incorporation of external sulfur to the  
577 magma as evidenced by the non-mantle-like S isotopic signatures of various sulfide deposits (e.g., Li  
578 et al. 2003; Ripley et al. 1999, 2002, 2005; Barnes et al. 2001; Duan et al. 2016).

579  
580 In this respect, the Kevitsa deposit is not an exception as almost all measured  $\delta^{34}\text{S}$  isotope values  
581 in both "barren" and mineralized samples in the Kevitsa ultramafic cumulates exceed +2 ‰ (Fig.  
582 5), suggesting incorporation of crustal sulfur into the Kevitsa magma. Both the regular and false  
583 ore type record heavy S isotope compositions ( $\delta^{34}\text{S}$  averaging +4.0 ‰ and +6.5 ‰, respectively)  
584 and hence external sulfur appears to have been important in the formation of these mineralization  
585 styles. For the Ni-PGE ore type, the average *in-situ*  $\delta^{34}\text{S}$  value ( $\sim +2.6$  ‰) is only slightly above the  
586 assumed mantle values. However, significantly heavier whole-rock  $\delta^{34}\text{S}$  values of +3.7 to +8.8 ‰  
587 have also been reported (Grinenko et al. 2003) and hence there might be a substantial internal  
588 isotopic variation in these ore zones.

589  
590 The obtained Sr and S isotope compositions do not show any mutual correlation. In particular,  
591 outside the ore domain, the Sr isotopic compositions remain relatively constant throughout the  
592 stratigraphy whereas S isotopes vary significantly (Figs. 6 and 7). As discussed above, this suggests  
593 bulk contamination of the silicate magma at depth, followed by selective assimilation of crustal  
594 sulfur during the transportation of the magma into the Kevitsa magma chamber. This is in

595 agreement with Grinenko et al. (2003) who reported heavy S isotopic compositions (+18 ‰ on  
596 average) for the immediate sedimentary rocks around the Kevitsa intrusion but noted marked  
597 decoupling in the C contents and  $\delta^{13}\text{C}$  values between the Kevitsa ores and sediments.  
598 Consequently, they concluded that these sedimentary rocks could not act as the main source of  
599 sulfur and hence sulfur assimilation took place at a deeper level in the crust.

600

601 As discussed above, it is proposed that the ore-bearing domain of the intrusion formed via  
602 multiple emplacements of variably contaminated silicate magma and sulfide liquid. Dynamic  
603 systems are generally considered favorable for generating economic sulfide deposits because  
604 sulfides are able to interact with, and collect metals from, a large volume of magma (e.g., Naldrett  
605 2011). The vigorous emplacement of the ore domain of the Kevitsa intrusion is further supported  
606 by the presence of numerous dunitic inclusions and komatiitic xenoliths, highlighting the capacity  
607 of intruding magmas to break off fragments from adjacent wall rocks. Recently, Luolavirta et al. (in  
608 press) proposed that the flow rate of the magma decreased due to the entrapment of a large  
609 number of inclusions, which aided settling of the sulfide droplets. An alternative view by Le  
610 Vaillant et al. (2017) proposes that the regular ores formed under high R factors in an expanded  
611 convecting magma chamber.

612

613 The origin of the Ni-PGE ore stays enigmatic. Yang et al. (2013b) suggested that assimilation of Ni-  
614 rich sulfides from komatiitic xenoliths enriched the magma in Ni, leading to the crystallization of  
615 Ni-enriched olivine found in the Ni-PGE ore and the formation of this ore type. However, they  
616 acknowledge that this model fails to explain the peculiar isotopic and REE characteristics of the  
617 ore type. The crustal-like isotope compositions of the host rocks to the Ni-PGE ore suggest

618 significant incorporation of crustal material in their genesis. This is in contrast with the primitive  
619 mineral compositions as well as the mineralogy of the host rocks (olivine-clinopyroxene cumulates  
620 with little or no orthopyroxene), which argue against any excessive role of crustal contamination.  
621 The contaminant for the Ni-PGE ore remains unidentified but was likely poor in silica and rich in  
622 calcium (and obviously Ni and PGEs) to aid crystallization of Ca-pyroxene rather than  
623 orthopyroxene. To generate the markedly distinct chemistry of the Ni-PGE ore in comparison to  
624 the regular and false ore, the magma(s) producing the Ni-PGE ore type probably interacted with  
625 different country rocks en route to the Kevitsa magma chamber.

626

627 The regular and false ore types record similar initial  $^{87}\text{Sr}/^{86}\text{Sr}$  (this study) and  $\epsilon_{\text{Nd}}$  (Huhma et al.  
628 2017 in review), REE characteristics (Hanski et al. 1997; Luolavirta et al. 2018) as well as a range in  
629 the compositions of ferromagnesian minerals. Hence, the parental silicate magmas for the two  
630 ore types were likely similar in composition. However, marked differences are observed in the S  
631 isotope compositions and metal contents in minerals and rocks, which are discussed in the  
632 following chapter.

633

#### 634 ***Previous sulfide segregation***

635

636 Olivine Fo-Ni trends are powerful in providing information about the evolution of the magmas,  
637 and in particular their sulfide saturation history (e.g., Li & Naldrett 1999; Li et al. 2004; Li et al.  
638 2007; Thakurta et al. 2008; Li et al. 2013). This is due to the fact that in the presence of sulfides, Ni  
639 as a highly chalcophile element ( $D^{\text{Ni}} \sim 500$ ; Peach et al. 1990) readily partitions into the sulfide

640 phase, resulting in Ni depletion in the magma and in the olivine crystallized from it. Also, sub-  
641 solidus Ni-Fe exchange reactions between olivine and coexisting sulfide may modify the primary  
642 olivine compositions, producing an inverse olivine Ni-Fo relationship as documented from several  
643 magmatic sulfide ore deposits, such as Jinchuan (Li et al. 2004), Noril'sk and Talnakh (Li et al. 2003)  
644 and Voisey's Bay (Li and Naldrett 1999).

645

646 In the Kevitsa olivine-pyroxene cumulates, sample-scale variations in the Fo content of olivine are  
647 modest (~1 %) but Ni in olivine may vary considerably, resulting in a wide scatter in a Fo vs. Ni plot  
648 (Fig. 9). The same is true for the olivine in the Central Dunite and dunite inclusions hosted by the  
649 Kevitsa intrusion (Luolavirta et al. in press). To model fractional crystallization, theoretical olivine  
650 Fo-Ni compositional trends were constructed using the PETROLOG software (Danyushevsky and  
651 Plechov 2011). The olivine-melt model (involving  $D^{\text{Ni}}$ ) from Herzberg and O'Hara (2002), the  
652 clinopyroxene-melt model from Danyushevsky (2001), and the clinopyroxene-melt  $D^{\text{Ni}}$  value of 3  
653 (Lindstrom and Weill 1978) were used. A Mg-rich picritic basalt (700 ppm Ni) from the Savukoski  
654 Group (Hanski et al. 2001a) is considered parental for the early-stage dunitic cumulates of the  
655 Kevitsa intrusive suite (Luolavirta et al. in press) and was used as the initial melt composition. The  
656 extremely nickeliferous olivines in the host rocks to the Ni-PGE ore were not considered in the  
657 calculation. Calculations were conducted under conditions of QFM+3 (three log units above the  
658 quartz-fayalite-magnetite  $f\text{O}_2$  buffer), QFM+2 and QFM-1.

659

660 The olivine in the host rocks to the false ore is clearly depleted in Ni with respect to the expected  
661 Ni contents due to fractional crystallization. Analyses of olivine enclosed in orthopyroxene  
662 oikocrysts (isolated from sulfides) in false ore samples record similarly low Ni abundances as  
663 olivine grains in contact with sulfides, indicating that the Ni-poor nature of olivine is not due to

664 late-stage equilibration with sulfides. Analogously, Yang et al. (2013b) did not observe any  
665 significant differences in the olivine Ni contents between olivine grains enclosed in clinopyroxene  
666 crystals and those in contact with sulfide minerals in the Ni-PGE ore type. Furthermore, we cannot  
667 identify any reverse trends in the Ni-Fo relationships that would suggest Fe-Ni exchange reactions.  
668 Therefore, the magma producing the olivine in false ore was most likely depleted in Ni due to an  
669 early attainment of sulfide saturation.

670

671 Given the compositional and isotopic similarity of the host rocks to the regular and false ores, the  
672 false ore could represent a differentiate after the formation of economic Ni-Cu ore, explaining the  
673 lack of metals in the former. The olivine data from Mutanen (1997) suggest that the forsterite  
674 content in olivine in false ore ( $\text{Fo}_{76-78.5}$ ) is overlapping but generally lower than that of the regular  
675 ore ( $\text{Fo}_{77-84}$ ). However, our new data on false ores expand the range of Fo in false ores to the  
676 range of regular ores. Also, the whole-rock compositions show no difference in the fractionation  
677 stage between these two ore types. It must be noted that the dense sulfide liquid may migrate  
678 within semi-consolidated cumulates, so that the observed assemblage of silicate and sulfide  
679 minerals in false ores do not necessarily represent interrelated accumulations. The different S  
680 isotope compositions of the regular and false ore, however, do not favor a simple genetic  
681 relationship of the two.

682

683 Due to the highly chalcophile nature of platinum group elements ( $D_{\text{PGE}}^{\text{sulfide-silicate}} \sim 20\,000$ ; e.g.,  
684 Fleet et al. 1991), the sulfide segregation history can also be assessed on the basis of the PGE  
685 contents and their relative abundances with respect to less chalcophile base metals (e.g., Maier et  
686 al. 1998; Barnes and Lightfoot 2005). In false ores, the Pt and Pd concentrations are low in



687 comparison to "barren" olivine pyroxenites (see Figs. 6 and 7) and fall close to or below the  
688 detection limit (10 ppb). Such low values may cause large errors in the calculated Cu/Pd or Pt/Pd  
689 ratios, for example, and hence, in this case, are not reliable for evaluating the possibility of an  
690 earlier sulfide-saturation event. Nevertheless, the markedly low concentrations of Pt and Pd in the  
691 false ores are well in line with the Ni-depleted nature of the olivine of the ore type, supporting an  
692 earlier sulfide segregation event. In this case, to produce metal-poor false ore, the magma must  
693 have reached sulfide saturation at two separate stages. The heavy sulfur isotope compositions  
694 measured for the false ore indicate a substantial proportion of crustal sulfur, which is consistent  
695 with multiple sulfur saturation events.

696

#### 697 ***Geological model***

698

699 The different Sr isotope profiles obtained from drill cores outside and within the ore domain  
700 cannot be unambiguously correlated. The Sr isotope ratios record a significant variability in the ore  
701 domain (see Fig. 8) whereas the surrounding rocks (see Figs. 6 and 7) show less radiogenic and  
702 more constant isotope compositions throughout the stratigraphy. Luolavirta et al. (2018)  
703 demonstrate that the lithological, whole-rock and mineral compositional variations in the  
704 stratigraphy are far more pronounced in the ultramafic rocks of the ore-bearing domain in  
705 comparison to the surrounding ultramafic rocks. This is well in line with the observed isotopic  
706 signatures. The isotopic and compositional differences can be interpreted to reflect distinct  
707 magmatic histories in the ore domain and its surroundings. To explain the spatial stratigraphic and  
708 compositional differences between different parts of the Kevitsa intrusion and the characteristic  
709 features of the ore types, the following integrated model is proposed for the magmatic evolution  
710 of the Kevitsa intrusive suite rocks and its Ni-Cu-(PGE) deposit (Fig. 10).

711 Stage 1. Intrusion(s) of picritic basalt magma formed olivine-chromite cumulates (Central Dunite)  
712 in the conduits (Luolavirta et al. in press).

713 Stage 2. Picritic basalt in a lower staging chamber differentiated to basaltic magma and underwent  
714 comprehensive country rock contamination. The basaltic magma intruded as a continuous, stable  
715 flow into the Kevitsa magma chamber and crystallized olivine-pyroxene cumulates. During their  
716 flow, the magma pulses assimilated sulfur from country rocks to variable degrees. When sulfur  
717 saturation was achieved, some of the sulfides were lodged probably in depressions within the  
718 conduit. Metal-depleted magma further gained sulfur from the country rocks and carried sulfide  
719 melt into the Kevitsa magma chamber precipitating metal-poor false ores (and contact type  
720 mineralization). Later crystal fractionation resulted in the formation of pyroxenites in a nearly  
721 closed system.

722 Stage 3. The Kevitsa magma chamber began to operate as a dynamic open system and magmas  
723 intruded into the hot interior of the Kevitsa intrusion. A number of stage-1 dunitic cumulates and  
724 xenoliths were further brecciated by these magmas and redistributed within the olivine-pyroxene  
725 cumulates. The origin of the zones of pOLWB remains ambiguous. They may either represent  
726 blocks of separate earlier sill(s) or autolith(s) of some marginal phase rocks formed at stage 2. A  
727 simple origin by crystal fractionation involving contamination cannot be ruled out either.

728 Ni-Cu-(PGE) ores formed by sulfur-saturated magma influxes, and were potentially upgraded by  
729 assimilation of the stage-2 proto-ore material in the conduit. Variable degrees of interaction of the  
730 new magma pulses with the early-formed sulfides could account for the wide variation in ore  
731 tenors found in the Kevitsa deposit. The magmas producing the Ni-PGE ore likely passed through  
732 distinct country rocks and followed a different route into the Kevitsa magma chamber. The

733 numerous dunite inclusions and mafic-ultramafic xenoliths within the deposit area may have  
734 reduced the flow rate of the magmas and aided concentration of the sulfides.

735

## 736 **Conclusions**

737

738 Both strontium and sulfur isotope compositions of the Kevitsa ultramafic cumulates deviate from  
739 those expected for purely mantle-derived magmas, indicating involvement of crustal material in  
740 their genesis. Variations in the Sr and S isotope data together with mineral and whole-rock  
741 chemistry demonstrate that the ore-bearing domain of the Kevitsa intrusion represents a dynamic  
742 site with multiple emplacements of variably contaminated silicate magma and sulfide liquid. In  
743 contrast, the rocks around the ore domain formed from a compositionally more uniform magma  
744 and underwent a less vigorous emplacement history. The sulfur in the Kevitsa ores is at least in  
745 part derived from crustal sources, with the external sulfur playing an important role in triggering  
746 sulfide saturation. The metal-depleted nature of the false ores and their host rocks indicate  
747 previous S-saturation and formation of proto-ores during an early stage of magma emplacement.  
748 Assimilation of the early formed sulfides may have upgraded the metal tenors of the Kevitsa Ni-  
749 Cu-(PGE) ore.

750

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762

763

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999

1000 **Table captions**

1001 **Table 1** Representative *in-situ* isotope data from plagioclase and sulfides from the Kevitsa  
1002 ultramafic zone.

1003 **Figure captions**

1004 **Fig. 1** Geologic map of the Kevitsa intrusion and location of the sampled drill cores and cross  
1005 sections of Fig. 2. Location of the Ni-Cu-(PGE) deposit is denoted by the outline of the open pit.

1006 **Fig. 2** a) S-N (A-A') and b) SE-NE (B-B') cross sections across the Kevitsa intrusion showing the  
1007 broad outlines of the Ni-Cu-(PGE) ore body (> 0.15 % Ni) and false ore bodies around the ore  
1008 domain (< 0.1 % Ni). For locations of the profiles, see Fig. 1.

1009 **Fig. 3** Photomicrographs of rocks from the Kevitsa ultramafic zone, showing laser spot positions as  
1010 red circles and  $Sr_{(i)}$  ratios with corresponding An contents in plagioclase. a) and b) olivine  
1011 pyroxenite, c) pyroxenite, d–f) plagioclase-bearing (ol) websterite.

1012 **Fig. 4**  $\epsilon_{Nd}$  vs.  $^{87}Sr/^{86}Sr_{(i)}$  diagram showing compositions of the Kevitsa false, regular and Ni-PGE ore  
1013 types.  $\epsilon_{Nd}$  data for Kevitsa ores from Huhma et al. (2017 in review) and Sr isotopic data from the



1014 present study. Field for Bushveld Bushveld Lower and Main zone and mantle-derived magma after  
1015 Maier et al. (2000) and for Archean tonalitic gneiss (Tojottamanselkä gneiss, 10 km north of  
1016 Kevitsa) after Hanski et al. (2001b) and Kröner et al. (1981).

1017 **Fig. 5** Sulfur isotopic compositions of Kevitsa ores and ultramafic rocks. Whole-rock data from  
1018 Grinenko et al. (2003). *In-situ* data from the present study.

1019 **Fig. 6** Stratigraphic variations of Sr isotope ratios and An contents of plagioclase and S isotope  
1020 compositions of sulfide in drill core KV-297 (outside the ore domain). The observed range in Sr  
1021 isotope compositions is depicted by the gray shaded column. Whole-rock Zr, La/Nb, S, Ni, Cu, Pt  
1022 and Pb contents are from the Kevitsa mine data base. Fo contents of olivine taken from Luolavirta  
1023 et al. (2018) and field for mantle-derived sulfur after Ripley and Li (2003). PX - Pyroxenite, OLPX -  
1024 olivine pyroxenite, Basal PX-GAB - Basal pyroxenite-gabbro. Po - pyrrhotite, Py - pyrite.

1025 **Fig. 7** Stratigraphic variations of Sr isotope ratios and An contents of plagioclase and S isotope  
1026 compositions of sulfide in drill core KV-280 (outside the ore domain). The observed range in Sr  
1027 isotope compositions is depicted by the gray shaded column. Whole-rock Zr, La/Nb, S, Ni, Cu, Pt  
1028 and Pb contents are from the Kevitsa mine data base. Fo contents of olivine taken from Luolavirta  
1029 et al. (2018) and field for mantle-derived sulfur after Ripley and Li (2003). PX - pyroxenite, OLPX -  
1030 olivine pyroxenite, Basal PX-GAB - Basal pyroxenite-gabbro.

1031 **Fig. 8** Stratigraphic variations of Sr isotope ratios and An contents of plagioclase and S isotope  
1032 compositions of sulfide in drill core KV-103. Gray shaded column represents the range in Sr isotope  
1033 compositions observed in drill cores KV-297 and KV-280 (outside the ore domain, Figs. 4 and 5).  
1034 Whole-rock Zr, La/Nb, S, Ni, Cu, Pt and Pb contents are from the Kevitsa mine data-base. Fo  
1035 contents of olivine taken from Luolavirta et al. (2017b). OLPX - olivine pyroxenite, \* - olivine  
1036 (clino)pyroxenite. pOLWB - plagioclase-bearing (olivine) websterite. Po - pyrrhotite, Py - pyrite.

1037 **Fig. 9** Plot of nickel vs. forsterite (Fo %) contents of olivine in dunitic cumulates and Kevitsa olivine-  
1038 pyroxene cumulates compared with theoretical olivine compositional trends calculated at QFM+3,  
1039 QFM+2 and QFM-1 for picritic-basalt parental magma. White dots in model curves refer to 10, 20,  
1040 30 and 40 percentages of fractional crystallization. Olivine data for dunitic cumulates taken from  
1041 Luolavirta et al. (in press) and for Kevitsa olivine-pyroxene cumulates from Luolavirta et al. (2018).

1042 **Fig. 10** Schematic illustration of the emplacement of the Kevitsa intrusive suite rocks and  
1043 formation of the Ni-Cu-(PGE) ore (modified after Luolavirta et al. 2017b).

1044

1045 **Electronic Supplementary Material 1**

1046 *In-situ*  $^{87}\text{Sr}/^{86}\text{Sr}$  ratios of the in-house standard determined during two sets of Sr isotope analyses.

1047 **Electronic Supplementary Material 2**

1048 *In-situ* Sr isotopic analyses of plagioclase from the Kevitsa intrusion.

1049 **Electronic Supplementary Material 3**

1050 Sample-scale correlations of *in-situ* determined  $^{87}\text{Sr}/^{86}\text{Sr}_{(i)}$  and anorthite contents of plagioclase.

1051 **Electronic Supplementary Material 4**

1052 *In-situ* S isotopic analyses of pyrrhotite and pyrite from the Kevitsa intrusion

1053

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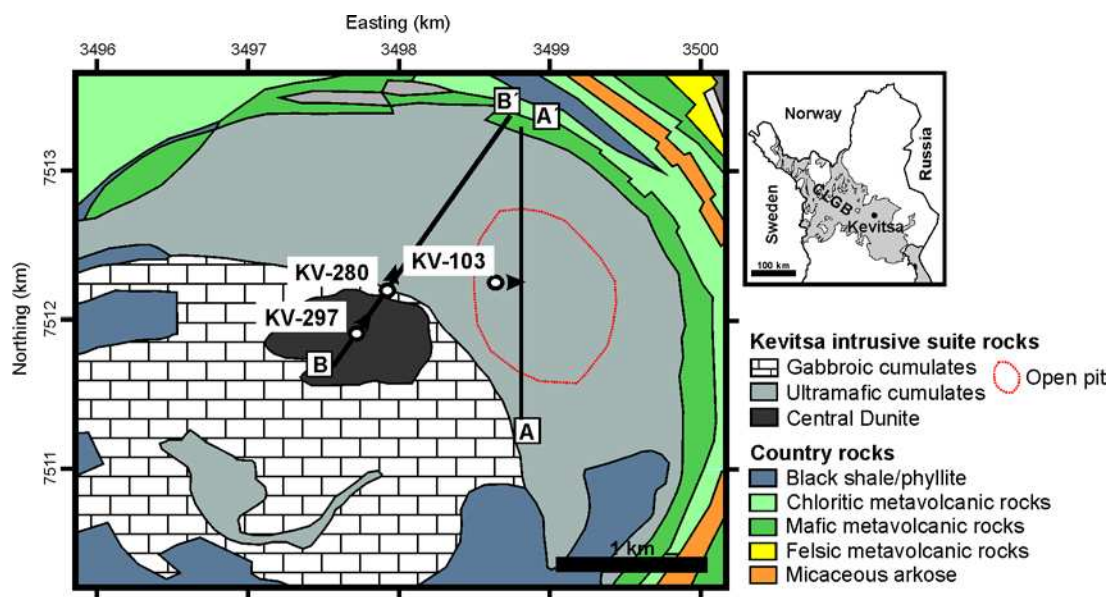
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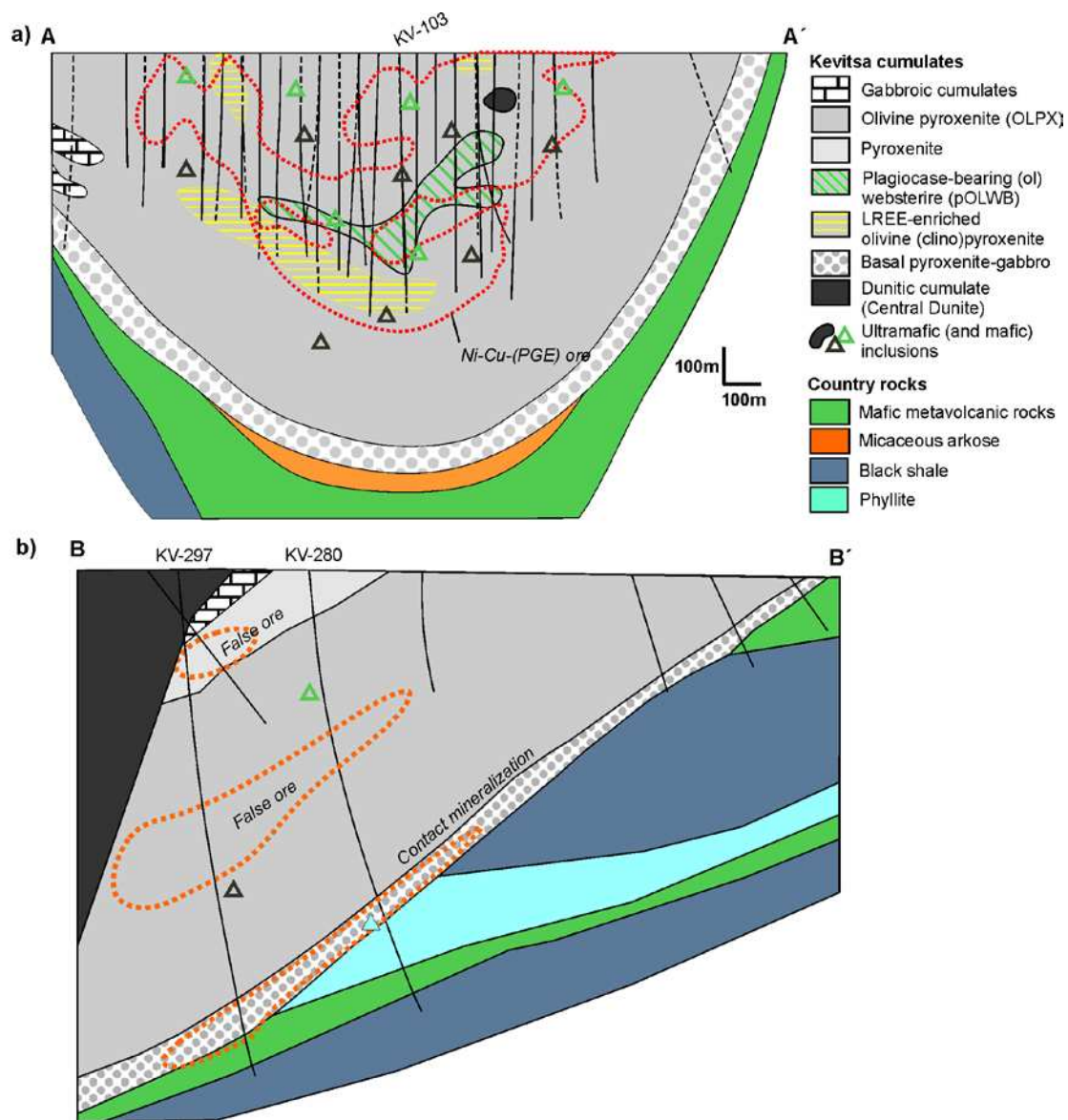
**Table 1** Representative *in-situ* isotope data from plagioclase and sulfides from the Kevitsa ultramafic zone

Sample	Lithology	<sup>87</sup> Rb/ <sup>86</sup> Sr	<sup>87</sup> Sr/ <sup>86</sup> Sr	<sup>87</sup> Sr/ <sup>86</sup> Sr <sub>(i)</sub>	2σ	An	δ <sup>34</sup> S‰	2σ	
KV280-74.45	PX	0.0049	0.7058	0.7056	0.00014	68.7			
		0.0061	0.7058	0.7056	0.00014	51.7			
		0.0071	0.7057	0.7055	0.00015	51.4			
KV297-391.95	OLPX	0.0020	0.7056	0.7056	0.00014	69.2	2.52	0.18	po
		0.0141	0.7054	0.7050	0.00015	70.4	3.04	0.18	po
KV280-574.95	OLPX (F)	0.0029	0.7051	0.7050	0.00012	68.8	7.60	0.20	po
		0.0026	0.7051	0.7050	0.00011	68.4	7.19	0.24	po
		0.0033	0.7054	0.7053	0.00013	72.8	8.00	0.23	po
		0.0027	0.7053	0.7052	0.00014	71.9			
KV297-1236.80	Basal PX	0.0032	0.7058	0.7057	0.00015	59.3	4.00	0.19	po
		0.0035	0.7051	0.7050	0.00014	60.6	3.57	0.21	po
KV103-258.08	OLPX	0.0033	0.7067	0.7066	0.00018	42.2	3.08	0.22	po
		0.0700	0.7091	0.7071	0.00029	42.4	3.35	0.24	po
KV103-303.49	OLPX	0.0130	0.7066	0.7062	0.00021	45.1	-0.42	0.15	py
		0.0116	0.7074	0.7071	0.00014	46.6	-0.33	0.26	po
		0.0082	0.7074	0.7072	0.00015	45.1	1.31	0.15	py
		0.0074	0.7069	0.7067	0.00021	45.5	1.55	0.23	po
KV103-421.10	OLPX (R)	0.0089	0.7064	0.7061	0.00023	46.2	4.13	0.24	po
		0.0111	0.7058	0.7055	0.00024	47.8	5.47	0.26	po
		0.0127	0.7063	0.7060	0.00022	44.5	3.75	0.20	po
		0.0069	0.7065	0.7063	0.00021	44.1	3.79	0.20	po
KV103-489.30	pOLWB	0.0106	0.7066	0.7063	0.00023	42.7 core	7.63	0.24	po
		0.0114	0.7078	0.7075	0.00024	42.7 core	6.98	0.19	po
		0.0103	0.7068	0.7065	0.00028	45.3 rim	6.99	0.20	po
		0.0127	0.7076	0.7072	0.00027	41.8 rim	5.84	0.23	po
KV103-671.18	OLPX* (N)	0.0095	0.7112	0.7109	0.00024	46.6	3.95	0.31	po
		0.0136	0.7107	0.7103	0.00030	47.6	1.33	0.24	po
		0.0112	0.7111	0.7108	0.00027	49.5	2.24	0.25	po

OLPX - olivine pyroxenite, OLPX\* - olivine (clino)pyroxenite, pOLWB - plagioclase bearing (ol) websterite, PX - pyroxenite  
F, R and N indicate false, regular and Ni-PGE ore, respectively. An=anorthite, po=pyrrhotite, py=pyrite.  
Initial <sup>87</sup>Sr/<sup>86</sup>Sr ratio ±2σ calculated using an age of 2.058 Ga



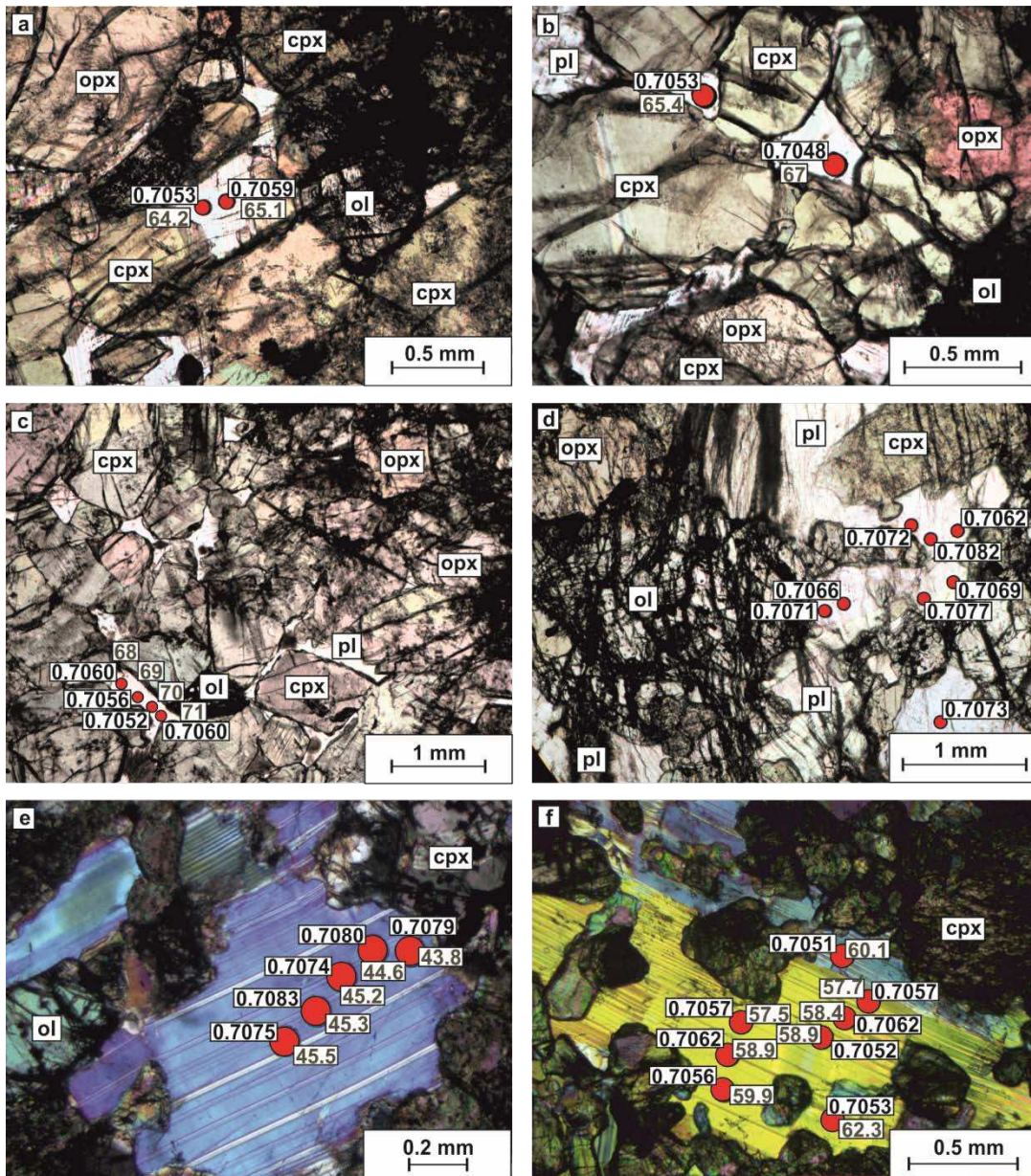
**Fig. 1**



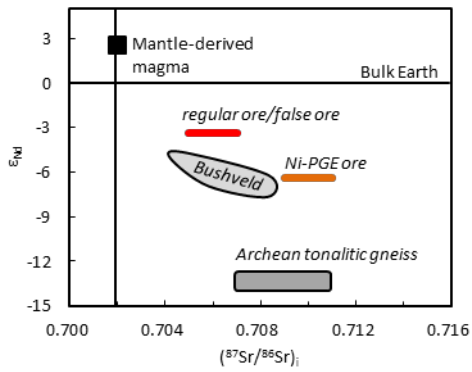
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1070 **Fig. 2**



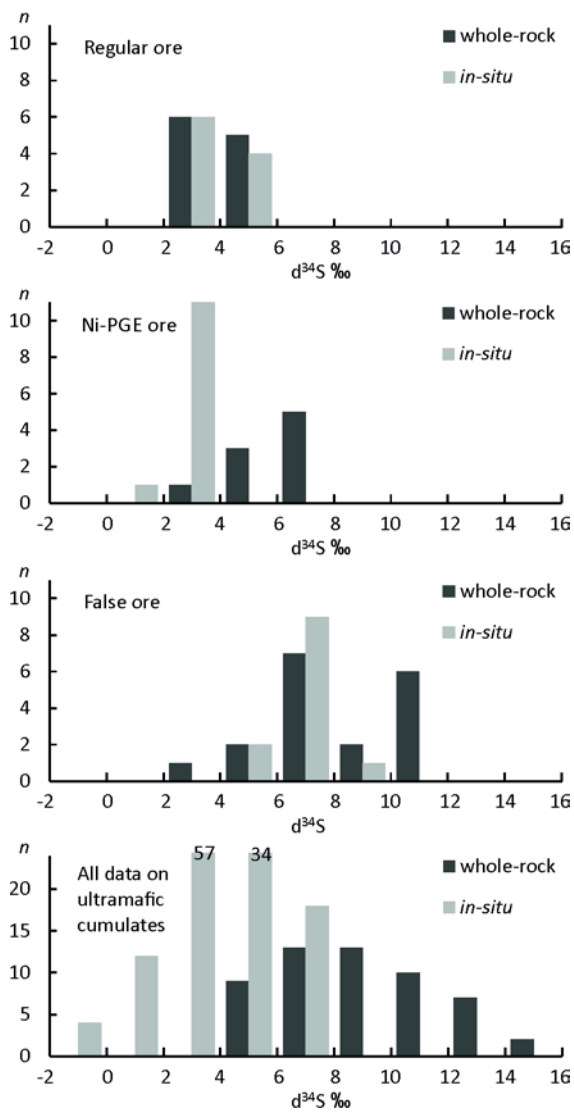


**Fig. 3**



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1076 **Fig. 4**

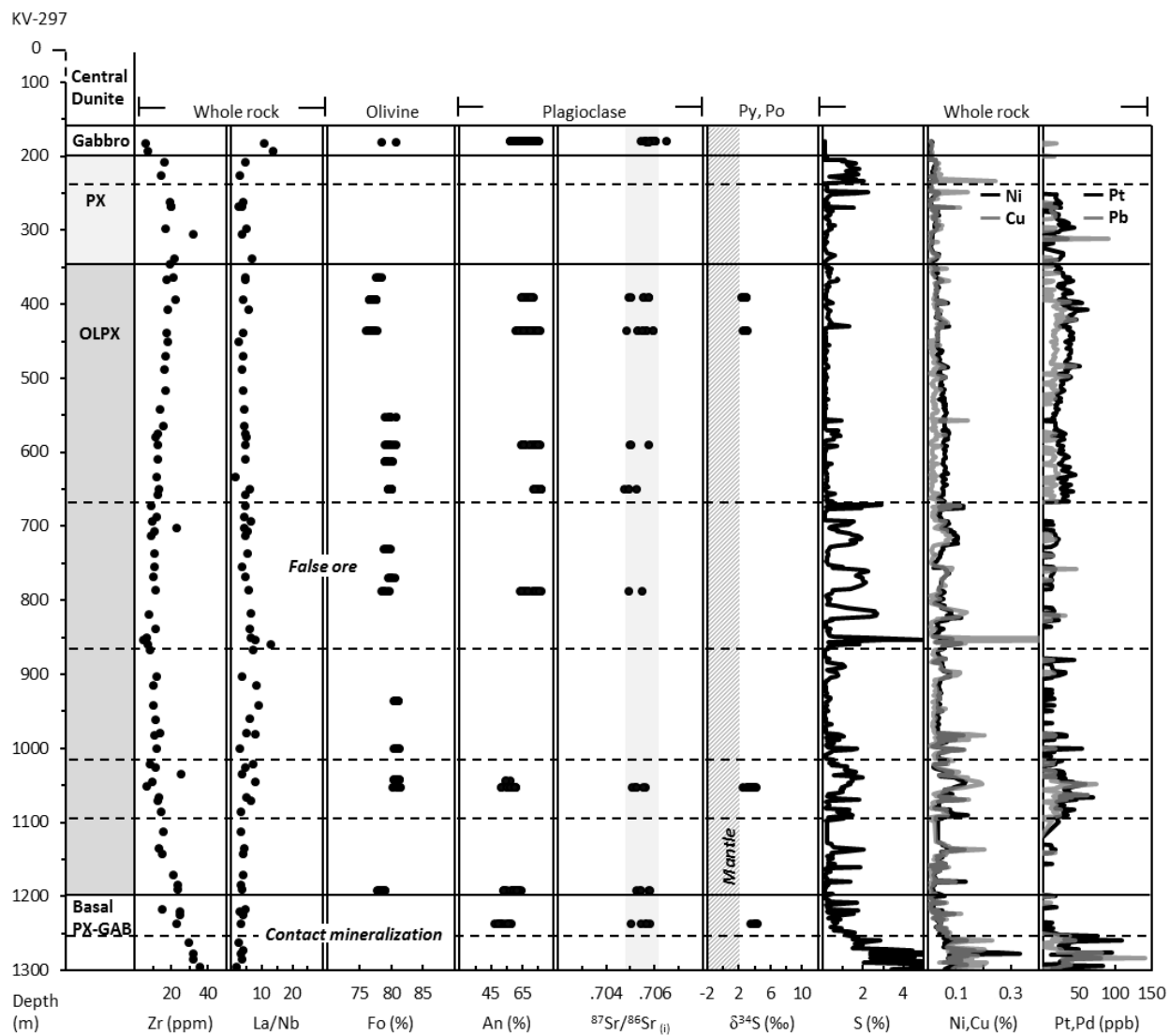


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1078 **Fig. 5**

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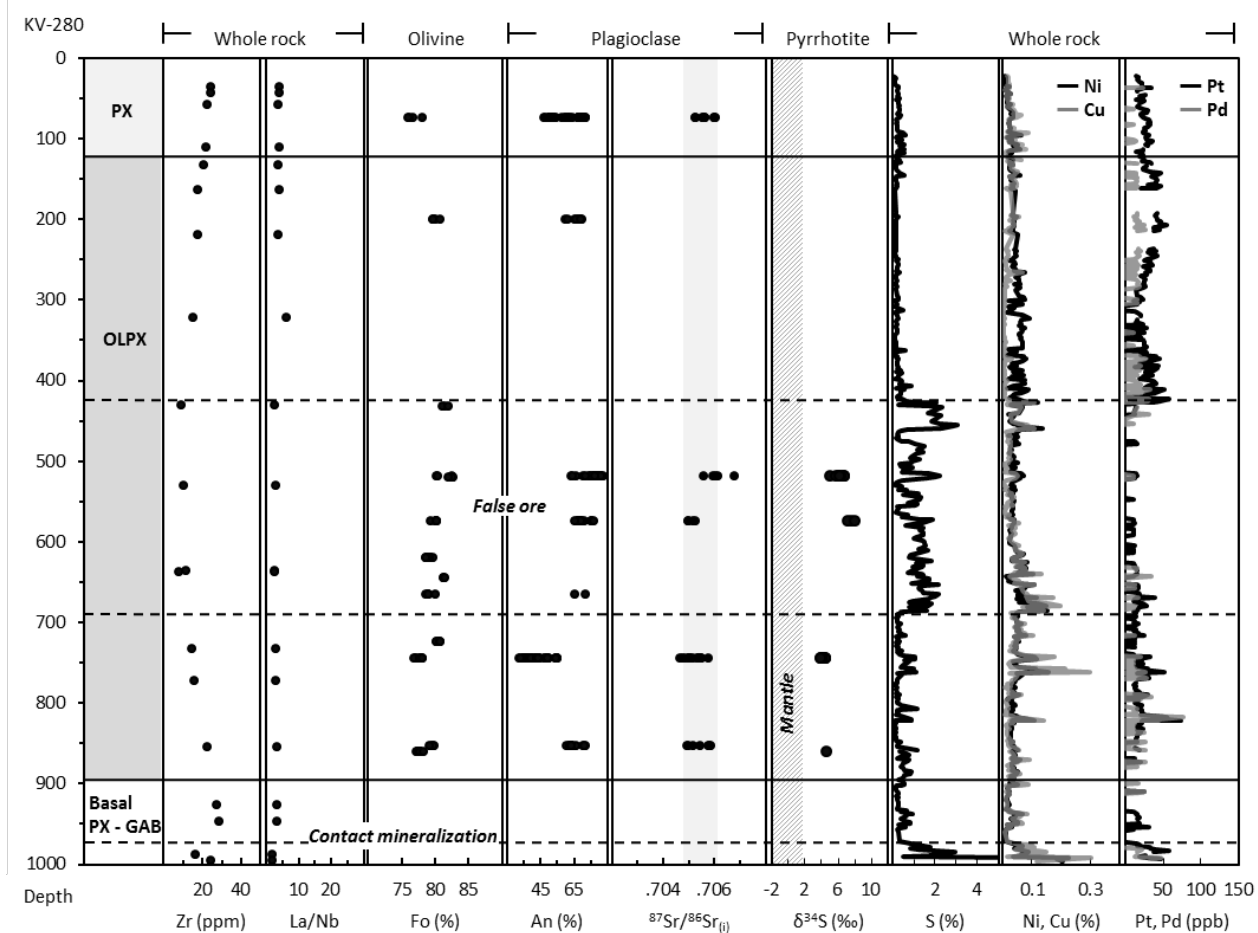
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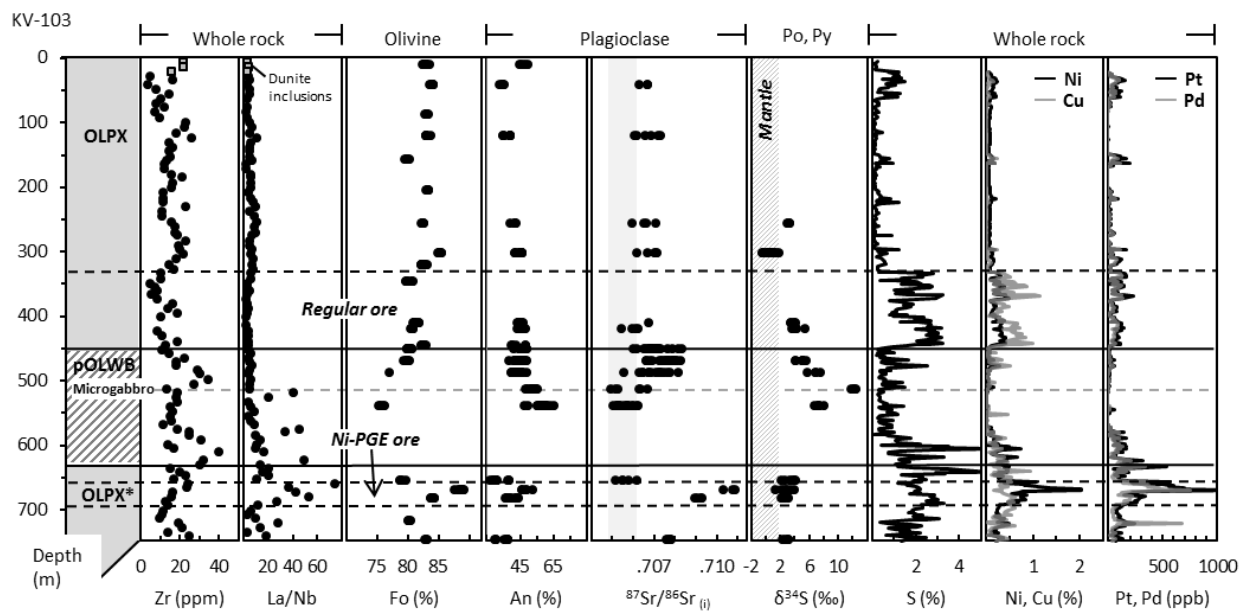
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1082 **Fig. 6**



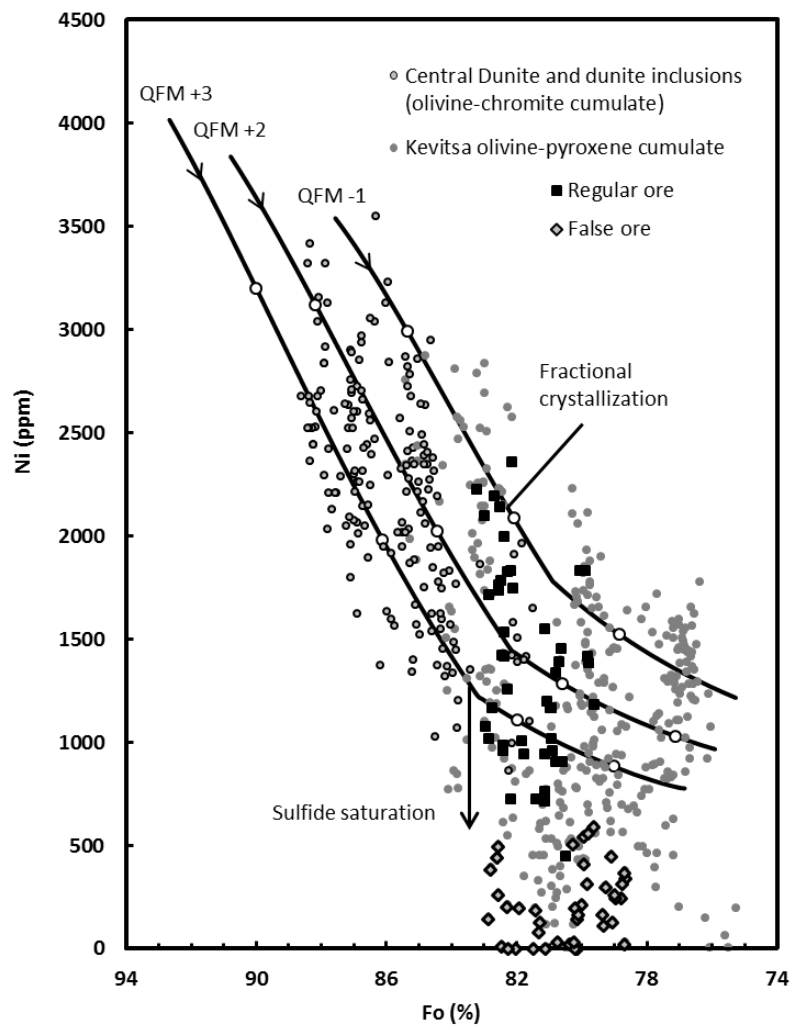


**Fig. 7**



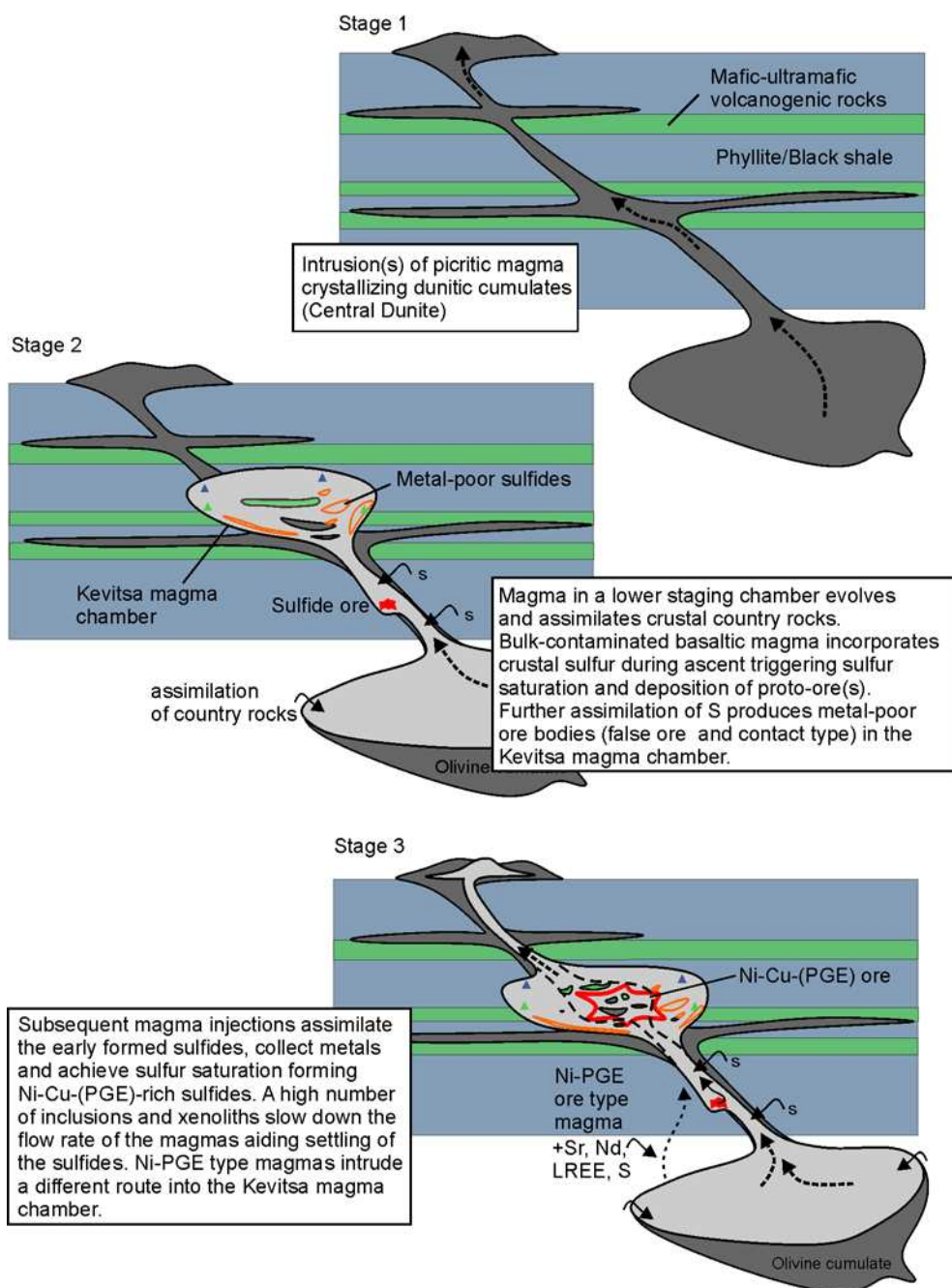
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1089 **Fig. 8**



1090

1091 **Fig. 9**



1092

1093 **Fig. 10**